

## Lecture 23, Nov 5, 2025

### Euler-Lagrange – Part 3

- *Kinetic energy* for a collection of point masses is  $T = \sum_{i=1}^N \frac{1}{2} m_i \|\dot{r}_i\|^2 = \sum_{i=1}^N \frac{1}{2} m_i \dot{r}_i^T \dot{r}_i$
- Notice  $\frac{d}{dt} \left( \frac{\partial T}{\partial \dot{q}_j} \right) = \sum_{i=1}^N \frac{d}{dt} \left( m_i \dot{r}_i^T \frac{\partial \dot{r}_i}{\partial \dot{q}_j} \right)$ 

$$= \sum_{i=1}^N \frac{d}{dt} \left( m_i \dot{r}_i^T \frac{\partial r_i}{\partial q_j} \right)$$
  - Note  $\dot{r}_i = \sum_{j=1}^n \frac{\partial r_i}{\partial q_j} \dot{q}_j$ , so  $\frac{\partial \dot{r}_i}{\partial \dot{q}_j} = \frac{\partial r_i}{\partial q_j}$
  - This is the first term in the summation for  $\sum_{i=1}^N m_i \ddot{r}_i^T \delta r_i$  from last lecture
  - Also,  $\frac{\partial T}{\partial q_j} = \sum_{i=1}^N m_i \dot{r}_i^T \frac{\partial \dot{r}_i}{\partial q_j}$  which is the second term
- Combining everything:  $\sum_{j=1}^n \left( \frac{d}{dt} \left( \frac{\partial T}{\partial \dot{q}_j} \right) - \frac{\partial T}{\partial q_j} \right) dq_j = \sum_{j=1}^n \varphi_j dq_j$ 
  - Since the  $dq_j$  are arbitrary, this means  $\frac{d}{dt} \left( \frac{\partial T}{\partial \dot{q}_j} \right) - \frac{\partial T}{\partial q_j} = \varphi_j, j = 1, \dots, n$
- For each force,  $f_i^l = f_i^g + f_i^a = -\nabla_{r_i} \mathcal{U}(r_1, \dots, r_N) + f_i^a = -\frac{\partial \mathcal{U}}{\partial r_i} + f_i^a$ , where the first term is the force caused by gravity
  - $\varphi_j = \sum_{i=1}^N (f_i^l)^T \frac{\partial r_i}{\partial q_j}$ 

$$= \sum_{i=1}^N -\frac{\partial \mathcal{U}}{\partial r_i} \frac{\partial r_i}{\partial q_j} + (f_i^a)^T \frac{\partial r_i}{\partial q_j}$$

$$= -\frac{\partial \mathcal{U}}{\partial q_j} + \tau_j$$
  - The term  $\tau_j$  captures all the other forces acting on the joint
- Finally, we get  $\frac{d}{dt} \left( \frac{\partial T}{\partial \dot{q}_j} \right) - \frac{\partial T}{\partial q_j} + \frac{\partial \mathcal{U}}{\partial q_j} = \tau_j$
- Let the *Lagrangian*  $\mathcal{L} = T - \mathcal{U}$ 
  - Note the potential energy is independent of  $\dot{q}_j$ , so  $\frac{\partial T}{\partial \dot{q}_j} = \frac{\partial \mathcal{L}}{\partial \dot{q}_j}$
- We arrive at the *Euler-Lagrange equations*:  $\frac{d}{dt} \left( \frac{\partial \mathcal{L}}{\partial \dot{q}_j} \right) - \frac{\partial \mathcal{L}}{\partial q_j} = \tau_j$

## Summary

The Euler-Lagrange equations are a set of  $n$  equations of motion for the system:

$$\frac{d}{dt} \left( \frac{\partial \mathcal{L}}{\partial \dot{q}_j} \right) - \frac{\partial \mathcal{L}}{\partial q_j} = \tau_j$$

where the Lagrangian is  $\mathcal{L} = T - \mathcal{U}$ ,  $n$  is the number of degrees of freedom ( $n = 3N - l$  for  $l$  constraints and  $N$  particles in 3 dimensions),  $q$  are  $n$  generalized coordinates which parametrize the set of allowed states,  $\tau_j$  are the generalized forces:

$$\tau_j = \sum_{i=1}^N (f_i^l)^T \frac{\partial r_i}{\partial q_j}$$

where  $f^l$  are the applied forces.

A set of  $l$  independent holonomic constraints are expressed as

$$g(r_1, \dots, r_N) = 0 \in \mathbb{R}^l \quad \text{rank} \left( \frac{\partial g}{\partial r} \right) = l$$

For constraints to be satisfied, the virtual displacements  $\delta r = [\delta r_1^T \quad \dots \quad \delta r_N^T]^T \in \mathbb{R}^{3N}$  must satisfy  $\frac{\partial g}{\partial r} \delta r = 0$ .