Lecture 28 (2-11), Mar 23, 2023

Ideal Gas (Sackur-Tetrode Formula)

• For a given U, how many microstates are there?

•
$$U = \sum_{i=1}^{N} \frac{\vec{p}_i^2}{2m} \implies \sum_{i=1}^{N} \vec{p}_i^2 = \sum_{i=1}^{N} p_{xi}^2 + p_{yi}^2 + p_{zi}^2 = 2mU$$

- How many ways can be distribute this momentum to get the same U?
- We want the sum of the squares of 3N numbers to equal 2mU
- This is a sphere that lives in 3N-dimensional space, or S^{3N-1} in \mathbb{R}^{3N} , with radius $r^2 = 2mU \implies r = \sqrt{2mU}$
- p_x, p_y, p_z are all quantized in units of $\frac{\pi\hbar}{L}$ (where $L^3 = V$), like in the case of the infinite square well; these points form a grid in space, the sphere will hit some of these grid points, and each hit is a microstate
 - We only need the sphere to be close to these grid points
 - We expect that the number of hits is proportional to the area of the sphere, proportional to r^{3N-1} , with r defined by the energy
- The area of a 3N 1-dimensional sphere in 3N dimensions can be shown to be $\frac{2\pi^{\frac{3N}{2}}}{\Gamma\left(\frac{3N}{2}\right)}R^{3N-1}$

- If
$$\frac{3N}{2}$$
 is an integer this is $\frac{2\pi^{\frac{3N}{2}}}{(\frac{3N}{2}-1)!}$
- Therefore $\Omega(N,U) \sim \frac{2\pi^{\frac{3}{2}}N}{(\frac{3}{2}N-1)!} \left(\sqrt{2mU}\right)^{3N-1}$, but we have to fix it first:

* The dimensions aren't consistent: to fix it, we divide $\sqrt{2mU}$ by a factor of $\frac{\pi\hbar}{L}$ first to make it unitless and make the units match

- This basically converts the p_i to n_i
- * The n_i are always positive, so the p_i must all be positive; this means we have to reduce the surface area by a factor of 2 for every axis
 - e.g. for a circle, if we restrict it to x, y > 0, we have to divide by a factor of 4; for a sphere, restricting it to the first quadrant divides by a factor of 8
- * The particles should be indistinguishable according to quantum mechanics, which means that if we swap the momenta of two particles, it stays in the same microstate
 - There are N! ways to permute the momenta, which all lead to identical microstates, so we have to reduce Ω by a factor of N!
 - This is called the Gibbs factor, which he derived before QM N^{3N-1}

$$-\Omega(N,U) = \frac{2\pi^{\frac{3}{2}}N}{N!\left(\frac{3}{2}N-1\right)!2^{3N}} \left(\frac{\sqrt{2mU}L}{\pi\hbar}\right)$$

- Now to find the entropy $S=k\ln\Omega$

- First use Stirling's approximation,
$$N! = \left(\frac{N}{e}\right)^N$$
 and ignore the -1
- Substitute $L = V^{\frac{1}{3}}$

$$-\Omega(N,U) = \frac{e^N}{N^N} \frac{2}{(2^3)^N} \frac{2(\pi^{\frac{3}{2}})^N (e^{\frac{3}{2}})^N}{(3^{\frac{3}{2}})^N (N^{\frac{3}{2}})^N} \frac{V^N}{(\pi\hbar)^{3N}} \left((2mU)^{\frac{3}{2}}\right)^N$$

$$= \dots$$
 I gave up ...

$$-S = k \ln \Omega$$
$$= k \ln 2 + k N \ln \left(\left(\left(\frac{4\pi mU}{N} \right)^{\frac{3}{2}} \frac{1}{(2\pi\hbar)^3} \frac{V}{N} \right) + \frac{5}{2} \right)$$
$$= k N \ln \left(\left(\left(\frac{4\pi mU}{N} \right)^{\frac{3}{2}} \frac{1}{(2\pi\hbar)^3} \frac{V}{N} \right) + \frac{5}{2} \right)$$

- * This is known as the Sackur-Tetrode formula for the entropy of an ideal gas Taking $\frac{1}{T} = \frac{\partial S}{\partial U} = Nk\frac{3}{2}\frac{1}{U} \implies U = \frac{3}{2}kNT$ which is the same as the one we got from equipartition
- What can we learn from this formula?
 - -S is on the same order as kN
 - For all the extensive quantities, e.g. N, U, V, if we double them, the entropy also doubles
 - -S increases with U, and levels off this is a normal system, the temperature increases with energy and we have a positive heat capacity
 - If we lower U enough, then eventually the stuff inside the log will become less than 1, and we'll get a negative entropy, which is impossible – therefore this formula does not work for gases with very low energy